Phenomenological formulation of relativistic spin hydrodynamics

Hidetoshi TAYA (Fudan University)

based on arXiv: 1901.06615 [hep-th]

in collaboration with

K. Hattori (YITP), M. Hongo (RIKEN), X.-G. Huang (Fudan), M. Matsuo (UCAS)

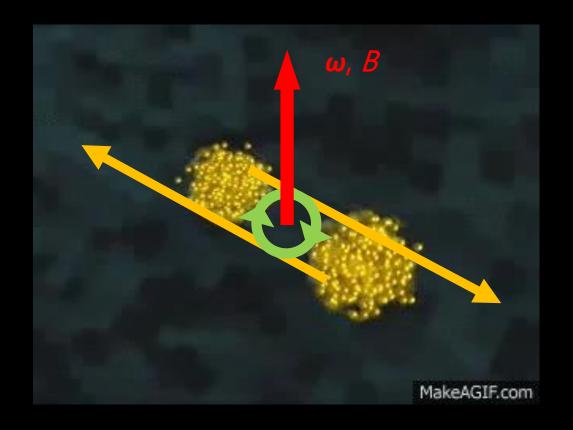
Ultra-relativistic heavy ion collisions



Aim: study quark-gluon plasma (QGP)

Found: QGP behaves like a perfect liquid and hydrodynamics works so well

Another interesting physics: Largest ω and B



Question: What happens under huge ω and/or B?

Specifically, any changes to QGP properties?

Naïve expectation: QGP is polarized

✓ Magnetic field B effect

Zeeman splitting (Landau quantization)

$$E \rightarrow E - s \cdot qB$$

charge dependent spin polarization

\checkmark Rotation ω effect

Bernett effect

$$E \rightarrow E - s \cdot \omega$$



→ charge *in* dependent spin polarization

Experimental fact

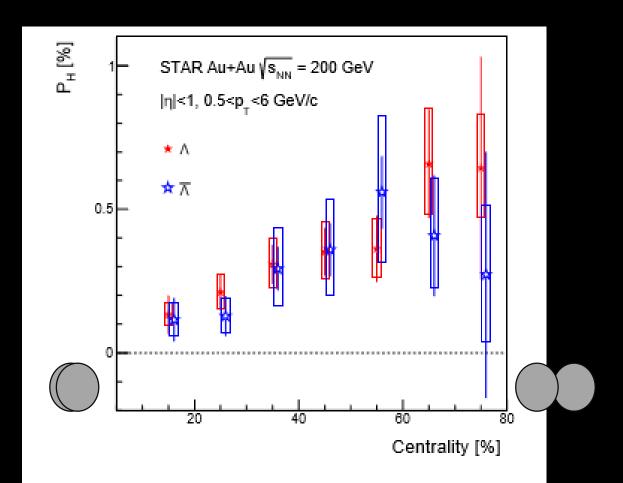


FIG. 5. Λ ($\bar{\Lambda}$) polarization as a function of the collision centrality in Au+Au collisions at $\sqrt{s_{NN}}=200$ GeV. Open boxes and vertical lines show systematic and statistical uncertainties. The data points for $\bar{\Lambda}$ are slightly shifted for visibility.

STAR (2018) See also talk by T.Niida

How about theory?

Hydrodynamics for spin polarized QGP?

→ Far from complete

Hydrodynamics for spin polarized QGP

✓ "Hydro simulations" exist, but...

usual hydro (i.e., hydro w/o spin) is solved

- (1) Compute velocity gradient at freezeout and define thermal vorticity $\widetilde{\omega}^{\mu\nu} \equiv \partial^{\mu}(u^{\nu}/T) \partial^{\nu}(u^{\mu}/T)$ Becattini, Chandra, Del Zanna, Grossi (2013)
- (2) Use Cooper-Frye formula with spin $f(E) \to f(E s \cdot \omega)$, where ω is spin chemical potential ($\neq \widetilde{\omega}$ in general) Becattini, Florkowski, Speranza (2018)
- (3) Assume $\omega = \tilde{\omega}$ (true only for global equilibrium)
- (4) Get spin-dependent hadron yield

✓ Formulation of relativistic spin hydrodynamics is still developing

Current status of formulation of spin hydro

✓ Non-relativistic case

e.g. Eringen (1998); Lukaszewicz (1999)

Already established (e.g. micropolar fluid)

- applied to pheno. and is successful
- e.g. spintronics: Takahashi et al. (2015)
- spin must be dissipative because of mutual conversion between spin and orbital angular momentum

✓ Relativistic case

Some trials exist, but

- only for "ideal" fluid (no dissipative corrections)
- some claim spin should be conserved

Purpose of this talk

- ✓ Formulate relativistic spin hydrodynamics with 1st order dissipative corrections for the first time
- ✓ Clarify spin should be dissipative

Outline

- 1. Introduction
- 2. Formulation based on an entropy-current analysis
- 3. Linear mode analysis
- 4. Summary

Outline

- 1. Introduction
- 2. Formulation based on an entropy-current analysis
- 3. Linear mode analysis
- 4. Summary

Introduction to hydrodynamics w/o spin (1/3)

Hydrodynamics is a low energy effective theory that describes spacetime evolution of IR modes (hydro modes)

textbook by Landau, Lifshitz

✓ Phenomenological formulation (EFT construction)

Step 1: Write down the conservation law: $0 = \partial_{\mu} T^{\mu\nu}$ 4 eqs

Step 2: Express $T^{\mu\nu}$ i.t.o hydro variables (constitutive relation)

- define hydro variables: $\{\beta, u^{\mu}\}$ $(u^2 = -1)$ $(u^2 = -1)$ $(u^2 = -1)$ $(u^2 = -1)$ $(u^2 = -1)$
- write down all the possible tensor structures of $T^{\mu\nu}$

$$\begin{split} T^{\mu\nu} &= f_1(\beta)g^{\mu\nu} + f_2(\beta)u^{\mu}u^{\nu} \\ &+ f_3(\beta)\epsilon^{\mu\nu\rho\sigma}\partial_{\rho}u_{\sigma} + f_4(\beta)\partial^{\mu}u^{\nu} + f_5(\beta)\partial^{\nu}u^{\mu} \\ &+ f_6(\beta)g^{\mu\nu}\partial^{\rho}u_{\rho} + f_7(\beta)u^{\mu}u^{\nu}\partial^{\rho}u_{\rho} + f_8(\beta)u^{\mu}\partial_{\mu}u^{\nu} + \dots + O(\partial^2) \end{split}$$

- simplify the tensor structures by (<u>assumptions</u> in hydro)
 - (1) symmetry
 - (2) power counting → gradient expansion
 - (3) other physical requirements → thermodynamics (see next slide)
- ✓ Hydrodynamic eq. = conservation law + constitutive relation

Introduction to hydrodynamics w/o spin (2/3)

✓ Constraints by thermodynamics

Expand $T^{\mu\nu}$ i.t.o derivatives

$$T^{\mu\nu} = T^{\mu\nu}_{(0)} + T^{\mu\nu}_{(1)} + O(\partial^2)$$

In static equilibrium $T^{\mu\nu} \to T^{\mu\nu}_{(0)} = (e, p, p, p)$, so that

$$T_{(0)}^{\mu\nu} = eu^{\mu}u^{\nu} + p(g^{\mu\nu} + u^{\mu}u^{\nu})$$

1st law of thermodynamics says

$$ds = \beta de$$
, $s = \beta (e + p)$

By using EoM $0 = \partial_{\mu} T^{\mu\nu}$, div. of entropy current $S^{\mu} = su^{\mu} + O(\partial)$ can be evaluated as

$$\partial_{\mu}S^{\mu} = -T^{\mu\nu}_{(1)}\partial_{\mu}(\beta u_{\nu}) + O(\partial^{3})$$

2st law of thermodynamics requires $\partial_{\mu}S^{\mu} \geq 0$, which is guaranteed if **RHS** is expressed as a semi-positive bilinear as

$$-T_{(1)}^{\mu\nu}\partial_{\mu}(\beta u_{\nu}) = \sum_{X_i \in T_{(1)}} \lambda_i X_i^{\mu\nu} X_{i\nu\mu} \ge 0 \text{ with } \lambda_i \ge 0$$

ex) heat current:
$$2h^{(\mu}u^{\nu)} \equiv h^{\mu}u^{\nu} + h^{\nu}u^{\mu} \in T^{\mu\nu}_{(1)} \ (u_{\mu}h^{\mu} = 0)$$

$$\Rightarrow T^{\mu\nu}_{(1)}\partial_{\mu}(\beta u_{\nu}) = -\beta h^{\mu}(\beta\partial_{\perp\mu}\beta^{-1} + u^{\nu}\partial_{\nu}u^{\mu}) \geq 0$$

$$\Rightarrow h^{\mu} = -\kappa(\beta\partial_{\perp\mu}\beta^{-1} + u^{\nu}\partial_{\nu}u^{\mu}) \text{ with } \kappa \geq 0$$

Introduction to hydrodynamics w/o spin (3/3)

✓ Constitutive relation up to 1st order w/o spin

$$T_{(0)}^{\mu\nu} = eu^{\mu}u^{\nu} + p(g^{\mu\nu} + u^{\mu}u^{\nu})$$

$$T_{(1)}^{\mu\nu} = -2\kappa \left(Du^{(\mu} + \beta\partial_{\perp}^{(\mu}\beta^{-1})u^{\nu}) - 2\eta\partial_{\perp}^{<\mu}u^{\nu>} - \zeta(\partial_{\mu}u^{\mu})\Delta^{\mu\nu}\right)$$
 heat current shear viscous effect bulk viscous effect

✓ Hydrodynamic equation w/o spin

Hydrodynamic eq. = conservation law + constitutive relation

Euler eq.

$$0 = \partial_{\mu} T^{\mu\nu}$$

$$T^{\mu\nu} = T^{\mu\nu}_{(0)}$$

Navier-Stokes eq.

$$0 = \partial_{\mu} T^{\mu\nu}$$

$$T^{\mu\nu} = T^{\mu\nu}_{(0)} + T^{\mu\nu}_{(1)}$$

:

Formulation of hydrodynamics with spin (1/4)

✓ Strategy is the same

Phenomenological formulation

<u>Step 1</u>: Write down the conservation law

<u>Step 2</u>: Construct a constitutive relation

- define hydro variables
- write down all the possible tensor structures
- simplify the tensor structures
 - (1) symmetry
 - (2) gradient expansion
 - (3) thermodynamics

Formulation of hydrodynamics with spin (2/4)

Step 1: Write down the conservation law

(1) energy conservation

$$0=\partial_{\mu}T^{\mu
u}$$
 (canonical)

(2) total angular momentum conservation

$$0 = \partial_{\mu} M^{\mu,\alpha\beta} \qquad \psi(x) \to S(\Lambda) \psi(\Lambda^{-1} x)$$

$$= \partial_{\mu} \left(L^{\mu,\alpha\beta} + \Sigma^{\mu,\alpha\beta} \right)$$

$$= \partial_{\mu} \left(x^{\alpha} T^{\mu\beta} - x^{\beta} T^{\mu\alpha} + \Sigma^{\mu,\alpha\beta} \right)$$

$$\partial_{\mu} \Sigma^{\mu,\alpha\beta} = T^{\alpha\beta} - T^{\beta\alpha}$$

- ✓ Spin is **not** conserved if (canonical) $T^{\mu\nu}$ has anti-symmetric part $T^{\mu\nu}_{(a)}$
- ✓ There's **no** a priori reason (canonical) $T^{\mu\nu}$ must be symmetric

Consequence

- (1) Spin must not be a hydro mode in a strict sense
- (2) Nevertheless, it behaves *like* a hydro mode if $T_{(a)}^{\mu\nu} \ll 1$
 - → inclusion of dissipative nature is important

Formulation of hydrodynamics with spin (3/4)

Step 2: Construct a constitutive relation for $T^{\mu\nu}$, $\Sigma^{\mu,\alpha\beta}$

(1) define hydro variables

4 + 6 = 10 DoGs = # of EoMs

Introduce spin chemical potential $\{\beta, u^{\mu}, \omega^{\mu\nu}\}$ with $\omega^{\mu\nu} = -\omega^{\nu\mu}$

 \checkmark { β , u^{μ} , $\omega^{\mu\nu}$ } are independent w/ each other at this stage ($\omega^{\mu\nu}$ \neq thermal vorticity)

(2) simplify the tensor structure by thermodynamics

Expand $T^{\mu\nu}$, $\Sigma^{\mu,\alpha\beta}$, i.t.o derivatives

$$T^{\mu\nu} = e u^{\mu} u^{\nu} + p (g^{\mu\nu} + u^{\mu} u^{\nu}) + T^{\mu\nu}_{(1)} + O(\partial^2), \quad \Sigma^{\mu,\alpha\beta} = u^{\mu} \sigma^{\alpha\beta} + O(\partial^1)$$

where I defined spin density $\sigma^{lphaeta}$

Generalizing 1st law of thermodynamics with spin as

$$ds = \beta(de - \omega_{\mu\nu}d\sigma^{\mu\nu}), \ \ s = \beta(e + p - \omega_{\mu\nu}\sigma^{\mu\nu})$$

With EoMs, div. of entropy current $S^{\mu} = su^{\mu} + O(\partial)$ can be evaluated as

$$\partial_{\mu}S^{\mu} = -T^{\mu\nu}_{(1s)} \frac{\partial_{\mu}(\beta u_{\nu}) + \partial_{\nu}(\beta u_{\mu})}{2} - T^{\mu\nu}_{(1a)} \left\{ \frac{\partial_{\mu}(\beta u_{\nu}) - \partial_{\nu}(\beta u_{\mu})}{2} - 2\beta \omega_{\mu\nu} \right\} + O(\partial^{3})$$

- ✓ In global equilibrium $\partial_{\mu}S^{\mu}=0$, so that $\omega=$ thermal vorticity.
- \checkmark 2nd law of thermodynamics $\partial_{\mu}S^{\mu} \geq 0$ gives strong constraint on $T_{(1)}^{\mu\nu}$

Formulation of hydrodynamics with spin (4/4)

✓ Constitutive relation for $T^{\mu\nu}$ up to 1st order with spin

$$T_{(0)}^{\mu\nu} = eu^{\mu}u^{\nu} + p(g^{\mu\nu} + u^{\mu}u^{\nu})$$

heat current shear viscous effect bulk viscous effect

$$T_{(1)}^{\mu\nu} = -2\kappa \left(Du^{(\mu} + \beta \partial_{\perp}^{(\mu} \beta^{-1}) u^{\nu)} - 2\eta \partial_{\perp}^{<\mu} u^{\nu>} - \zeta (\partial_{\mu} u^{\mu}) \Delta^{\mu\nu} \right)$$

$$-2\lambda \left(-Du^{[\mu}+\beta\partial_{\perp}^{[\mu}\beta^{-1}+4u_{\rho}\omega^{\rho[\mu}\right)u^{\nu]}-2\gamma \left(\partial_{\perp}^{[\mu}u^{\nu]}-2\Delta_{\rho}^{\mu}\Delta_{\lambda}^{\nu}\omega^{\rho\lambda}\right)$$

"boost heat current" "rotational (spinning) viscous effect"

NEW!

e.g. Eringen (1998); Lukaszewicz (1999)

- ✓ Relativistic generalization of a non-relativistic micropolar fluid
- ✓ "boost heat current" is a relativistic effect

✓ Hydrodynamics equation up to 1st order with spin

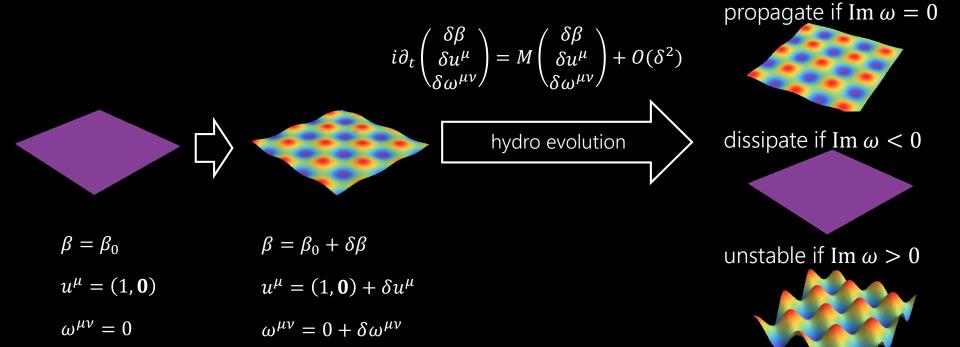
$$0 = \partial_{\mu} (T_{(0)}^{\mu\nu} + T_{(1)}^{\mu\nu} + O(\partial^{2})) \qquad \qquad \partial_{\mu} (u^{\mu} \sigma^{\alpha\beta}) = T_{(1)}^{\alpha\beta} - T_{(1)}^{\beta\alpha} + O(\partial^{2})$$

Outline

- 1. Introduction
- 2. Formulation based on an entropy-current analysis
- 3. Linear mode analysis
- 4. Summary

Linear mode analysis (1/2)

Setup: small perturbations on top of static equilibrium



Linear mode analysis (2/2)

✓ Hydro w/o spin $\{\beta, u^{\mu}\}$

4 modes

2 sound modes $\omega = \pm c_s k + O(k^2)$

2 shear modes $\omega = -i \frac{\eta k^2}{e+p} + O(k^4)$

where $c_{\rm s}^2 \equiv \partial p/\partial e^{\dagger}$

✓ Hydro with spin $\{\beta, u^{\mu}, \omega^{\mu\nu}\}$

4 modes

2 sound modes $\omega = \pm c_{\rm s}k + O(k^2)$

2 shear modes
$$\omega = -i \frac{\eta k^2}{e+p} + O(k^4)$$

+ 6 dissipative modes

3 "boost" modes $\omega = -2i\tau_{\rm b}^{-1} + O(k^2)$

3 "spin" modes $\omega = -2i\tau_s^{-1} + O(k^2)$

where
$$au_{\rm S}\equiv rac{\partial\sigma^{ij}/\partial\omega^{ij}}{4\gamma}$$
 , $au_{\rm b}\equiv rac{\partial\sigma^{i0}/\partial\omega^{i0}}{4\lambda}$

- ✓ We explicitly confirmed that spin is dissipative
- \checkmark The time-scale of the dissipation is controlled by the new viscous constants γ , λ

Outline

- 1. Introduction
- 2. Formulation based on an entropy-current analysis
- 3. Linear mode analysis

4. Summary

Summary

- ✓ Spin polarization of QGP is one of the hottest topics in HIC. But, its theory, in particular hydrodynamic framework, has not been developed well
- ✓ Relativistic spin hydrodynamics with 1st order dissipative corrections is formulated based on the phenomenological entropy-current analysis
- ✓ **Spin must be dissipative** because of the mutual conversion between the orbital angular momentum and spin
- ✓ Linear mode analysis of the spin hydrodynamic equation also suggests that spin must be dissipative

Outlook: extension to 2nd order, Kubo formula, application to cond-mat, numerical simulations, and start something new with you!