

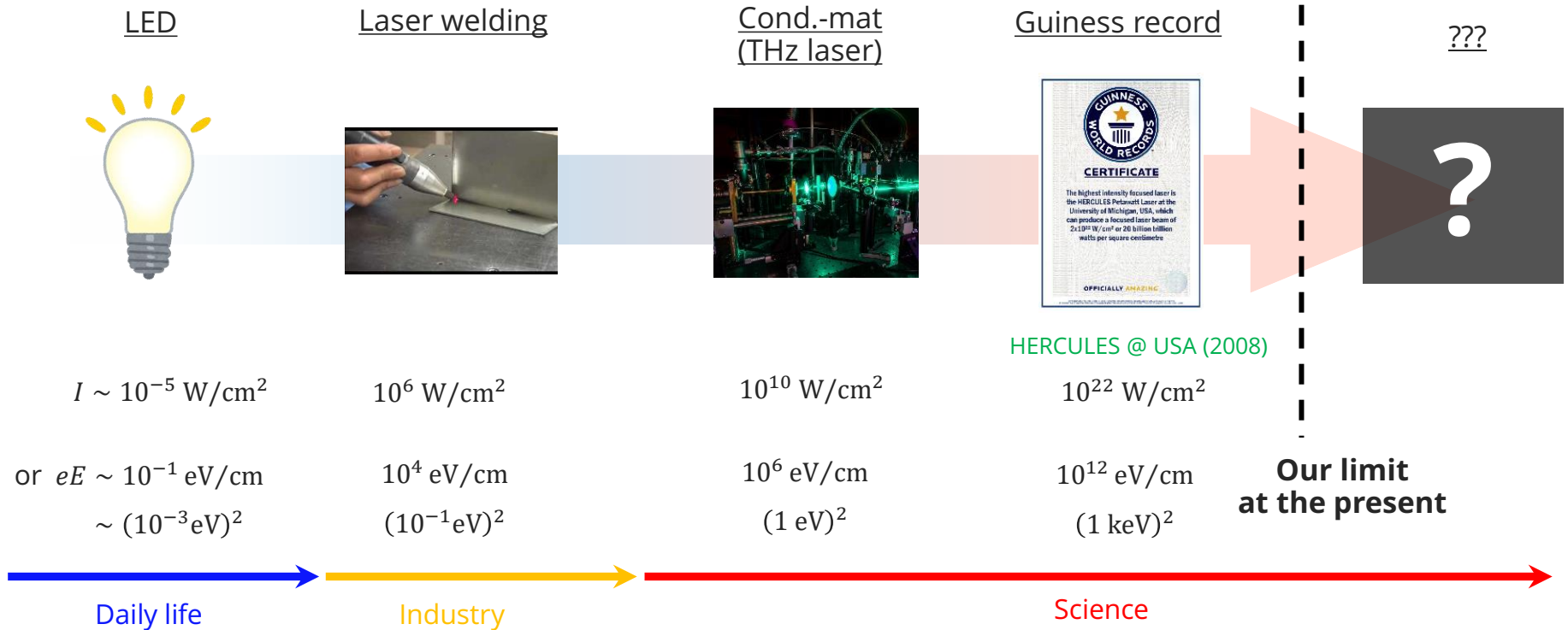
Introduction to strong-field physics

Hidetoshi Taya

Research area: Theoretical physics (particle/nuclear theory)

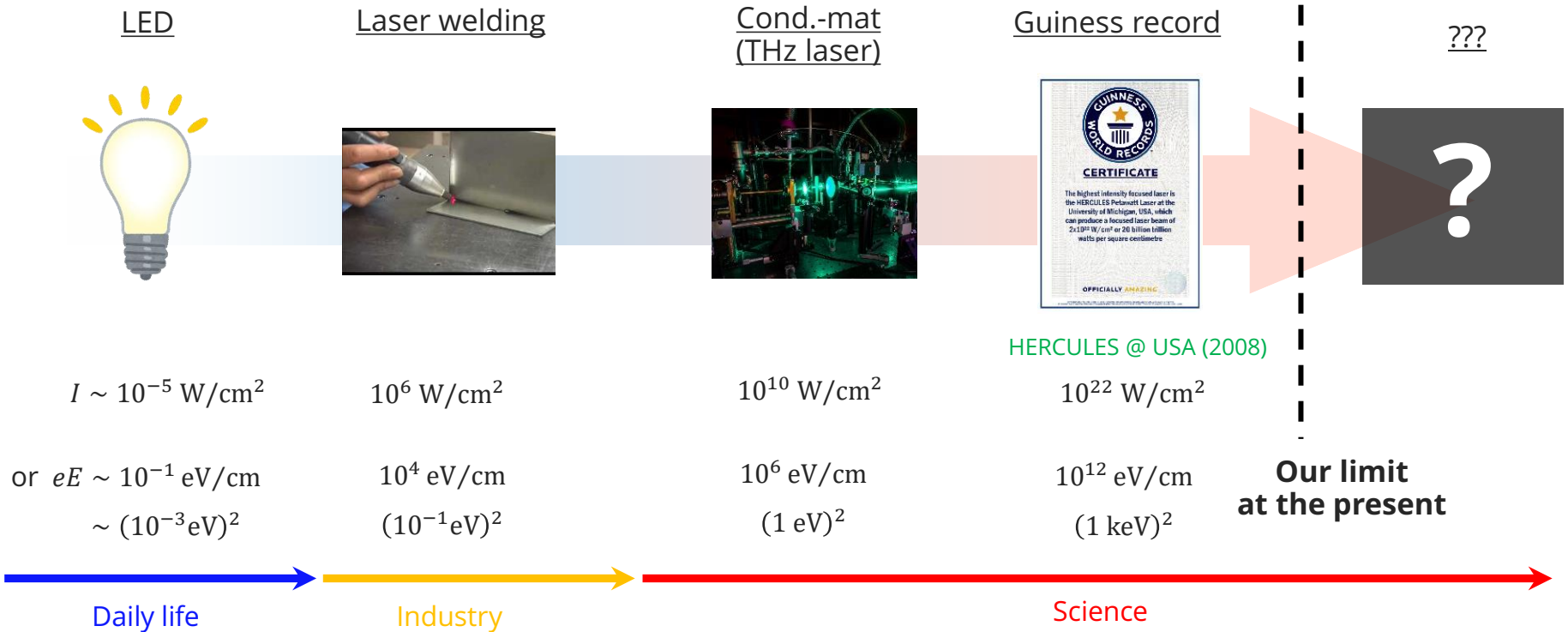
Today's talk

Q: What happens if we make light (or "field" in general) stronger and stronger ?



Today's talk

Q: What happens if we make light (or "field" in general) stronger and stronger ?



A: Our vacuum shall be destroyed and gives us nontrivial responses
 (\approx "something" even from "nothing")

\Rightarrow Purpose of this talk: Review such physics of strong light

Contents

1. Overview of strong-field physics

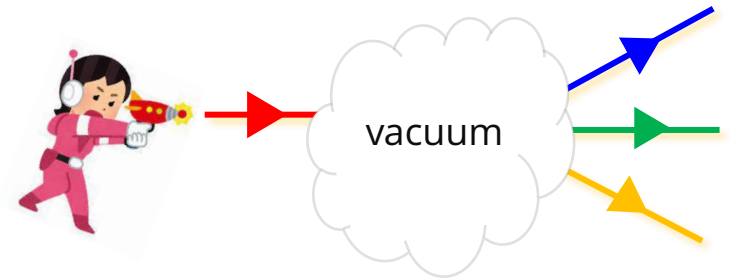
- becoming a hot topic in physics, thanks to the recent developments in exp.
- the Schwinger effect = a process to create “something” from “nothing”

2. Application of the Schwinger effect

[HT, Hongo, Ikeda (2021)]

⇒ High-harmonic generation from the vacuum

- first prediction of HHG from the vacuum
- a good example of “interdisciplinarity” of strong-field physics
- ⇒ your expertise should be useful in my area and vice versa !



3. Summary

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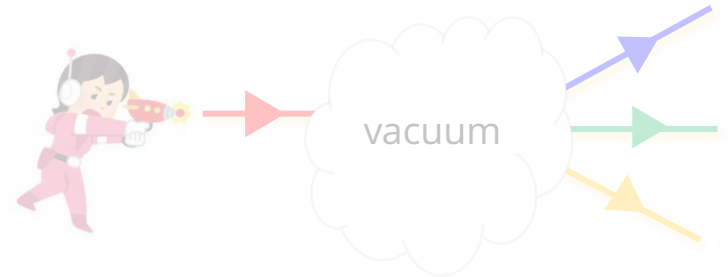
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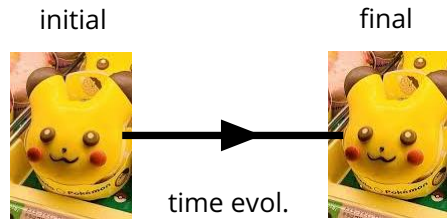
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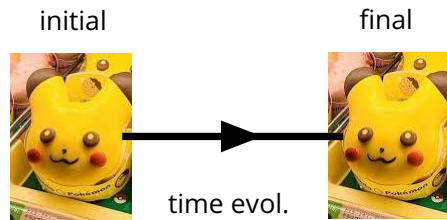
3. Summary

What if field becomes strong ?



No field

What if field becomes strong ?



No field

Weak field

Strong field



What if field becomes strong ?



No field

Weak field

Strong field

Only minor changes

⇒ Perturbative

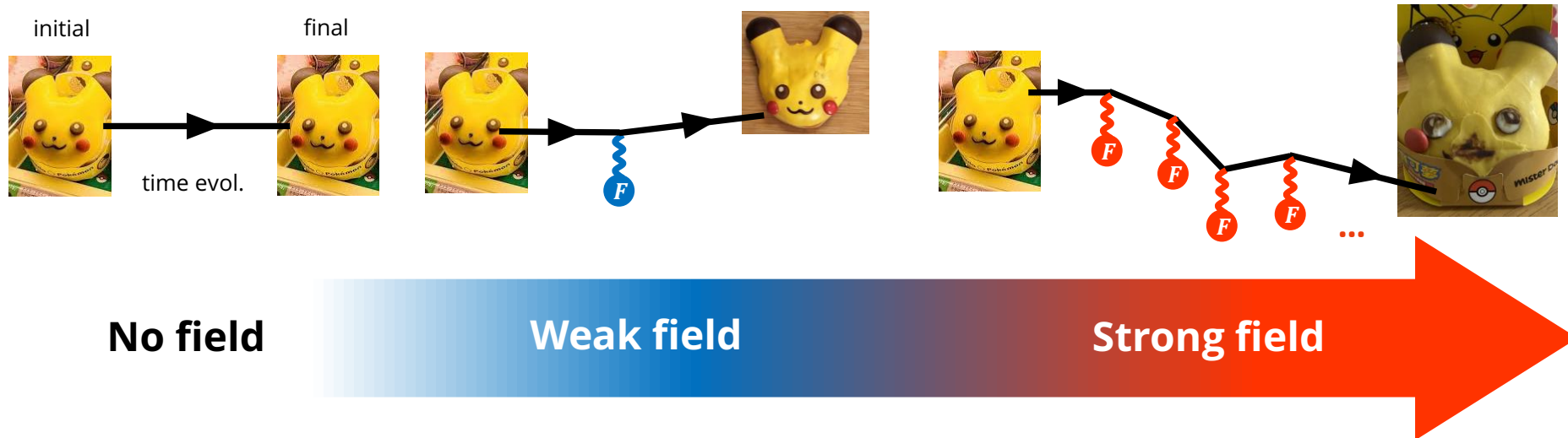
⇒ Very well understood
in both exp.& theor.

ex.) Electron (anomalous) magnetic moment $a := \frac{g-2}{2}$
≈ Electron energy shift in a weak magnetic field

$$a(\text{theor.}) = 1159652182.03 \dots \times 10^{-12}$$

$$a(\text{exp.}) = 1159652180.73 \dots \times 10^{-12} \quad [\text{Aoyama, Kinoshita, Nio (2017)}]$$

What if field becomes strong ?



Only minor changes

⇒ Perturbative

⇒ Very well understood
in both exp.& theor.

Big change !

⇒ Non-Perturbative

⇒ Not understood well

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If field becomes strong, physics becomes totally different & nontrivial

When is field “strong” ?

Strong-field condition:

To significantly modify the original system with typical energy Δ , the field must be more energetic than Δ

⇒ **Strong-field condition: $\Delta < (\text{energy scale of the field})$**

When is field “strong” ?

Strong-field condition:

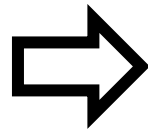
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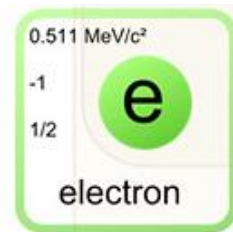
Estimate of the minimum field strength:

Standard Model

mass → charge → spin →	$\pm 2.3 \text{ MeV}/c^2$ 2/3 1/2 u up	$\pm 1.275 \text{ GeV}/c^2$ 2/3 1/2 c charm	$\pm 173.37 \text{ GeV}/c^2$ 2/3 1/2 t top	0 1 g gluon	$\pm 126 \text{ GeV}/c^2$ 0 0 H Higgs boson
	$\pm 4.8 \text{ MeV}/c^2$ -1/3 1/2 d down	$\pm 95 \text{ MeV}/c^2$ -1/3 1/2 s strange	$\pm 4.18 \text{ GeV}/c^2$ -1/3 1/2 b bottom	0 1 γ photon	
QUARKS	$0.511 \text{ MeV}/c^2$ -1 1/2 e electron	$105.7 \text{ MeV}/c^2$ -1 1/2 μ muon	$1.777 \text{ GeV}/c^2$ -1 1/2 τ tau	0 1 Z Z boson	
LEPTONS	$< 2.2 \text{ eV}/c^2$ 0 1/2 ν_e electron neutrino	$105.67 \text{ MeV}/c^2$ 0 1/2 ν_μ muon neutrino	$1.777 \text{ MeV}/c^2$ 0 1/2 ν_τ tau neutrino	$80.4 \text{ GeV}/c^2$ 1 1 W W boson	GAUGE BOSONS



The (matter) particle having the minimum energy



- Charged ⇒ Couples to electromagnetic (EM) field
- $eE \text{ MeV}^2 > m_e^2 = (0.511 \text{ MeV})^2 \approx 0(10^{28} \text{ W/cm}^2)$

When is field “strong” ?

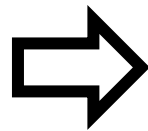
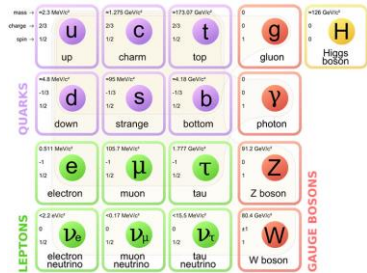
Strong-field condition:

To significantly modify the original system with typical energy Δ , the field must be more energetic than Δ

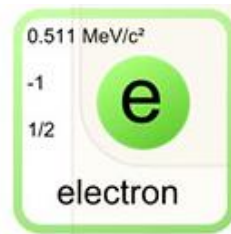
\Rightarrow Strong-field condition: $\Delta < (\text{energy scale of the field})$

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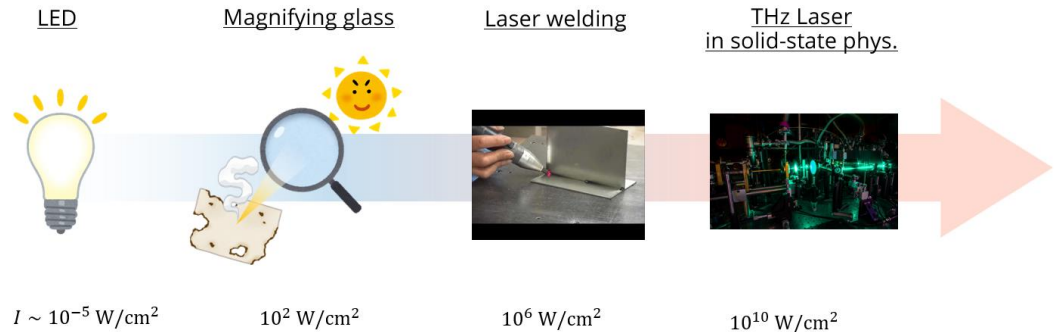
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∴ EXTREMELY strong

\Rightarrow Impossible to realize exp. in 20th century



Recent availability of strong EM fields

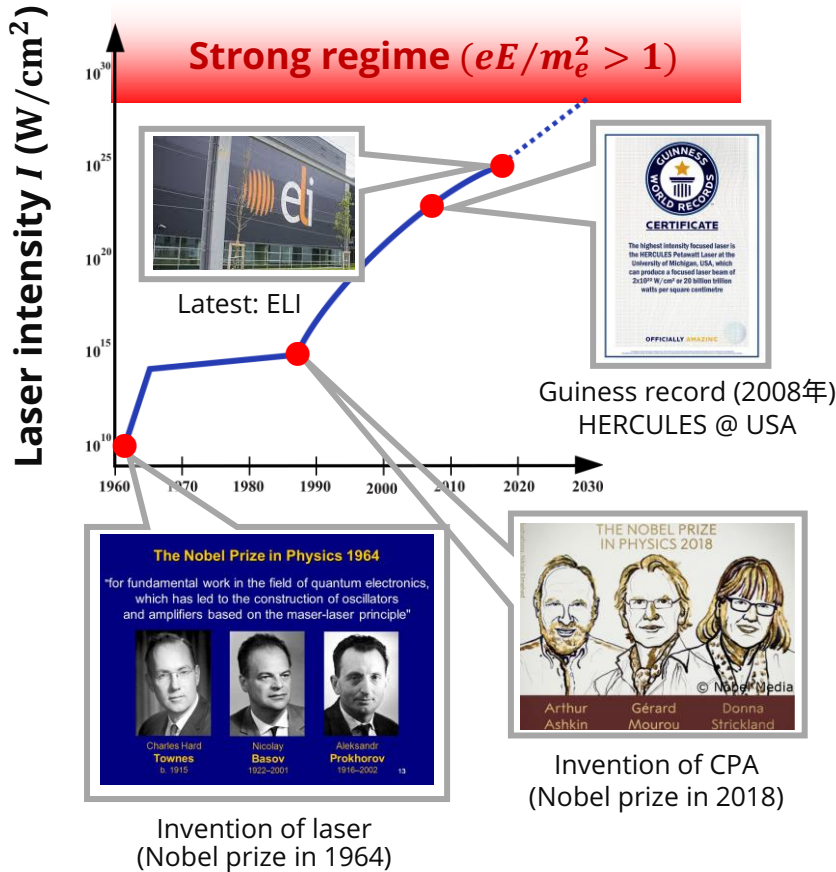
The situation changing: becoming able to create/observe strong light

∴ NOW is the BEST time to study strong-field physics

Recent availability of strong EM fields

The situation changing: becoming able to create/observe strong light

High-power laser



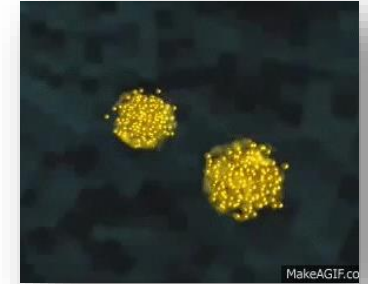
Extreme physical systems

• Heavy-ion collisions

RIC (2000~), LHC (2012~), FAIR/NICA/HIAF/J-Parc-HI/... (20XX~)

$$I \sim 10^{35} \text{ W/cm}^2$$

$$(eE, eB \sim m_\pi^2 \sim (140 \text{ MeV})^2)$$

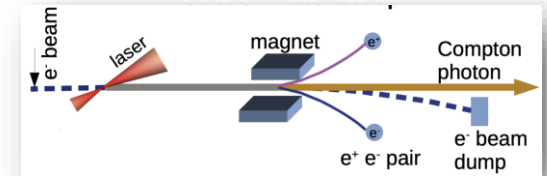


• Electron collider + Laser

Start soon: LUXE @ DESY, FACET-II @ SLAC

$$I \sim 10^{29} \text{ W/cm}^2$$

$$(eE, eB > m_e^2 \sim (1 \text{ MeV})^2)$$

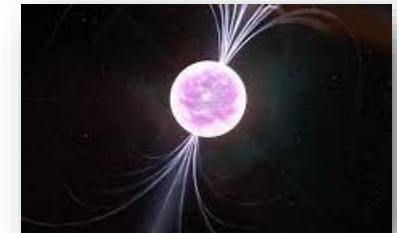


• Magnetar

Suzaku (2005~2015), NICER (2017~) XL-Calibur (2018~), IXPE (2021~), ...

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A bit more on heavy-ion collisions

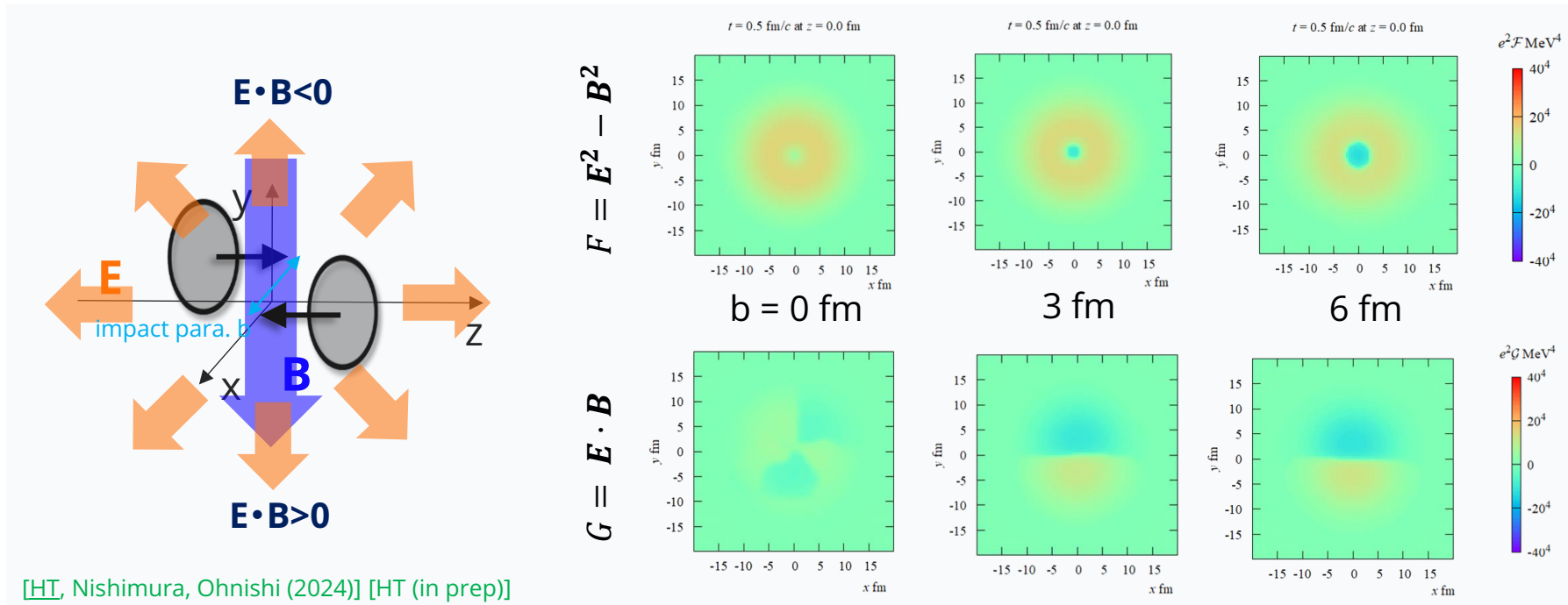
Low-energy ($\sqrt{s_{NN}} = 2 - 10$) HIC is interesting among other strong-field systems

- Sales point: The only system that has supercritical $F := E^2 - B^2 > 0$, $G := E \cdot B \neq 0$

	high-power laser	magnetar	High-energy HIC	Low-energy HIC
Field profile	(usually) $F = G = 0$	$F < 0, G = 0$	$F < 0, G = 0$	$F > 0, G \neq 0$
strength	subcritical	supercritical	far-supercritical	supercritical
lifetime	super-long	super-long	super-short	long

- Numerical estimation of the EM profile in low-energy HIC

@ $\sqrt{s_{NN}} = 5.2$ GeV

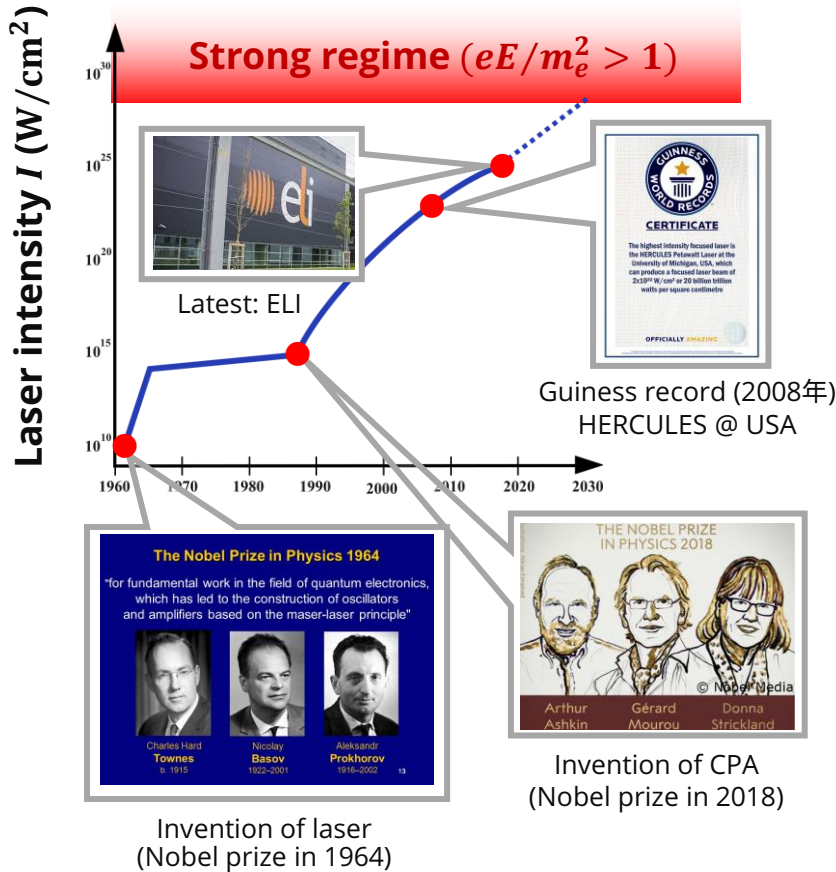


- So, I'm interested in this and wanna study this further:
 \Rightarrow Chiral XXX? Axion electrodynamics? Novel QCD phase? Let's discuss if interested 😊

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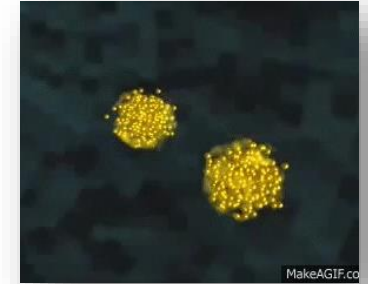
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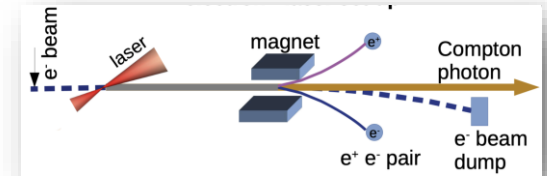


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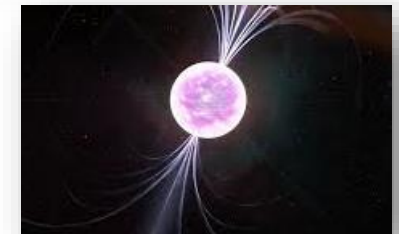


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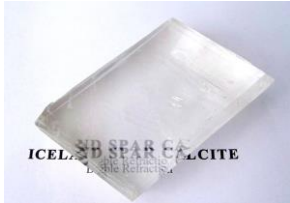
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What can happen with strong EM field ?

When $eE > m_e^2$, many non-trivial phenomena have been predicted to occur

Review: [Fedotov, Ilderton, Karbstein, King, Seipt, HT, Torgrimsson, Phys. Rept. (2023)]

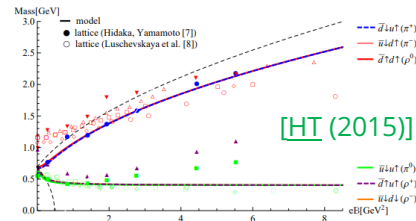
Birefringence of photon in vacuum



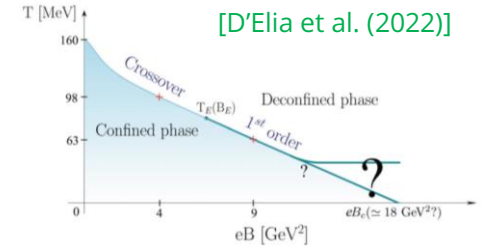
Photon splitting



Hadron mass change (for $eB > \Lambda_{\text{QCD}}^2$)

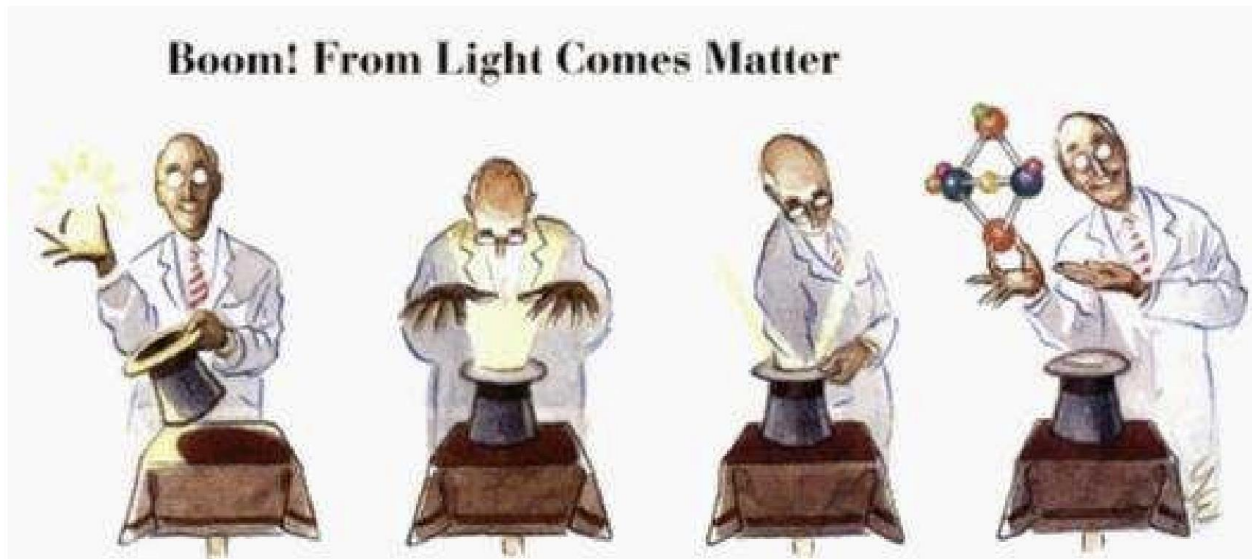


QCD phase diagram (for $eB > \Lambda_{\text{QCD}}^2$)



Among others, the (Sauter-)Schwinger effect is the most intriguing Sauter (1931), Schwinger (1951)

Particles are spontaneously produced from the vacuum (\approx "sthg" from "nothing")

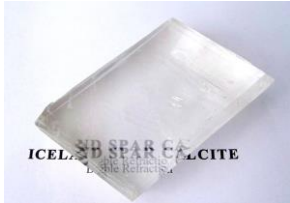


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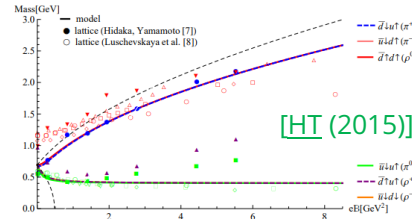
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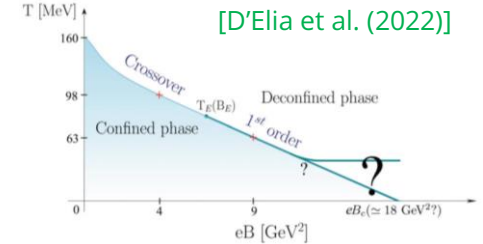
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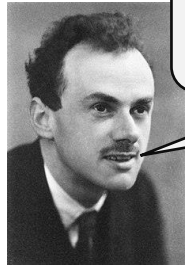
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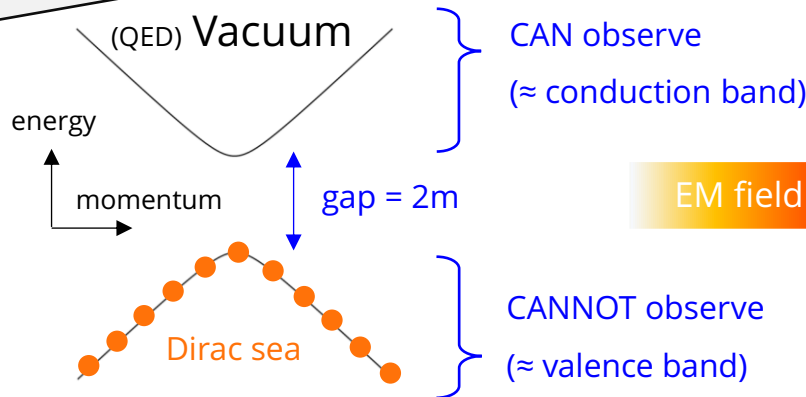
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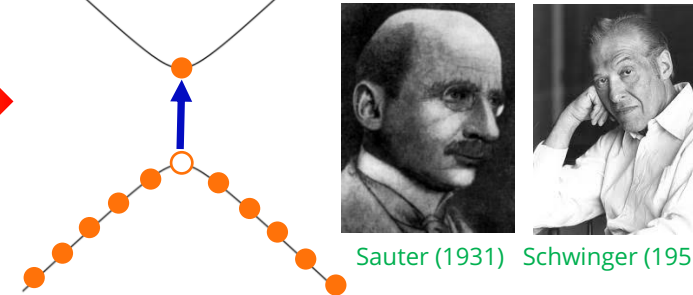
Dirac (1930)

Our vacuum is **not an empty space**, but has a structure **similar to semi-conductor**



Then, **real particles will be produced out of the vacuum** by strong enough field

Particle production !



Sauter (1931)



Schwinger (1951)

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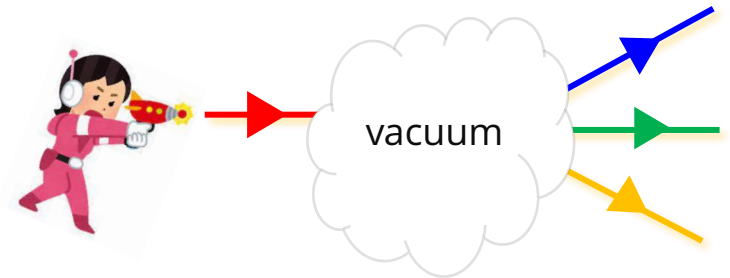
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Analogue of the Schwinger effect

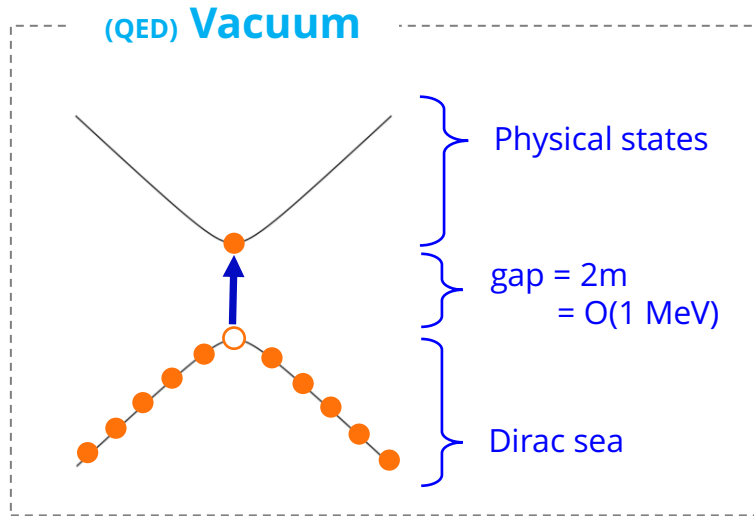
Analogous particle production occurs in various physical systems

- ex: Hawking radiation (strong gravity), dynamical Casimir effect (strong oscillation), (p)reheating in the early Universe (strong inflaton field), ...
- **Landau-Zener effect** in materials would be the most prominent [Landau \(1932\)](#), [Zener \(1932\)](#)

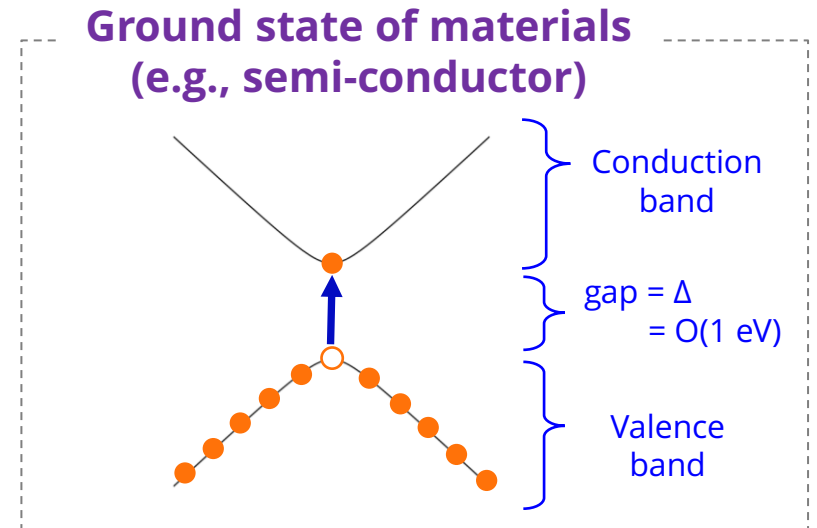
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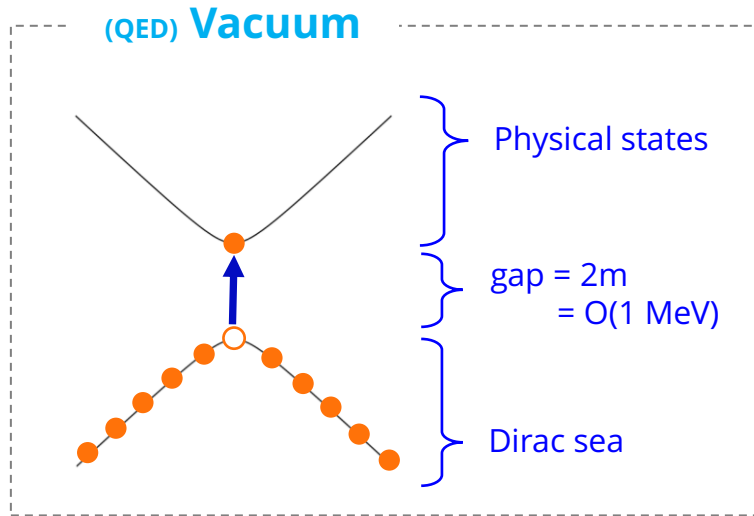


✓ The responses of the vacuum and ground state by an external EM field should be similar

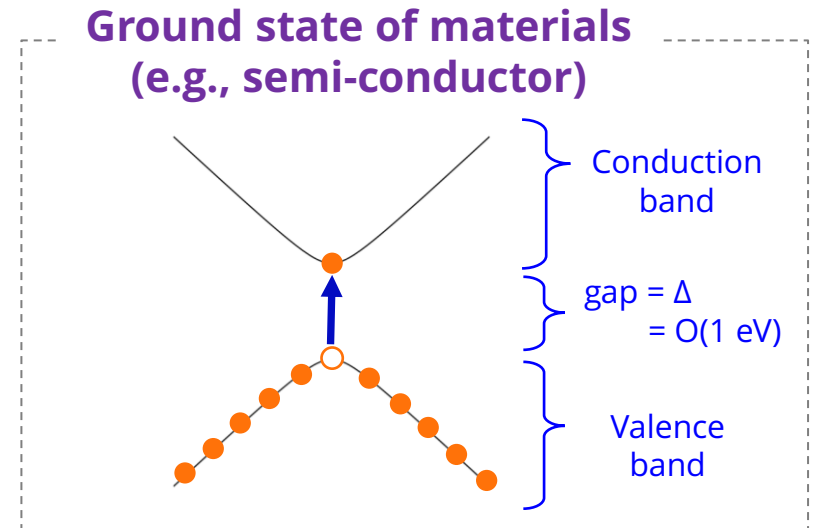
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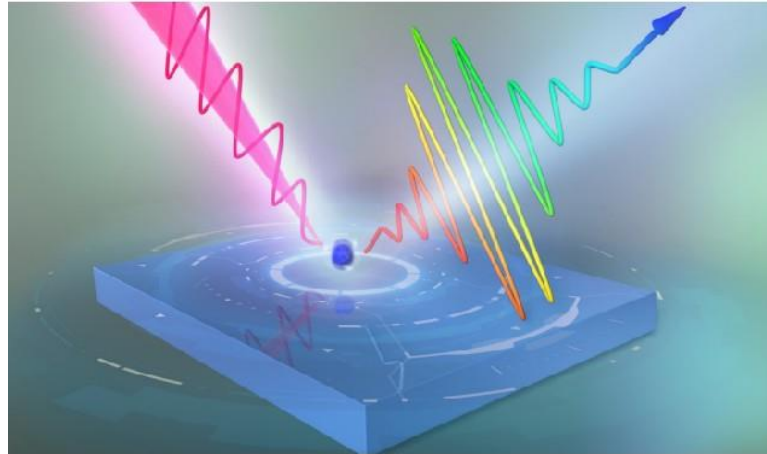
- ✓ The responses of the vacuum and ground state by an external EM field should be similar
- ✓ Indeed, the particle prediction rate and the excitation rate obey the same exponential formula

$$\Gamma_{\text{Schwinger}} = \# \exp\left[\# \frac{m^2}{eE}\right] \quad \Gamma_{\text{Landau-Zener}} = \# \exp\left[\# \frac{\Delta^2}{eE}\right]$$

- ✓ Schwinger effect has not been verified yet
BUT, Landau-Zener transition has been observed and utilized, e.g., to design devices

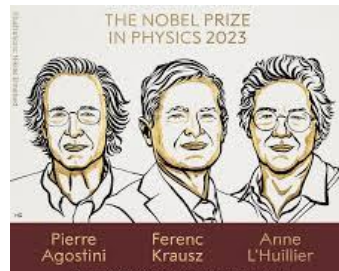
⇒ **Importing ideas from other areas of physics is quite to better understand the Schwinger effect (or strong-field physics in general)**

As such idea: High-harmonic generation from the vacuum

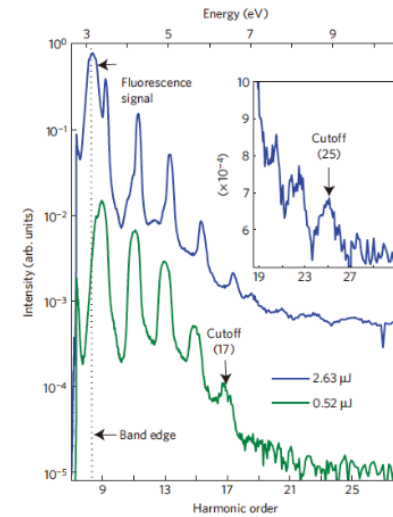
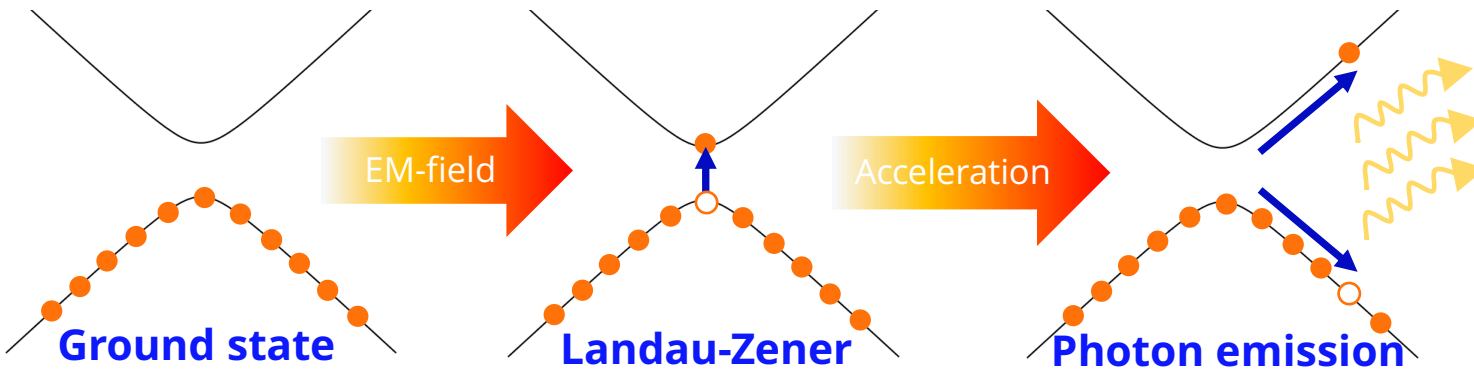


- **HHG developed in optics & solid-state physics**

- Observed with various materials: gas (1988), liquid (2014), solid (2011), ...
- Many applications: laser pointer, high-speed switching, ...
- Nobel prize in 2023



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[Ghimire et al., (2011)]

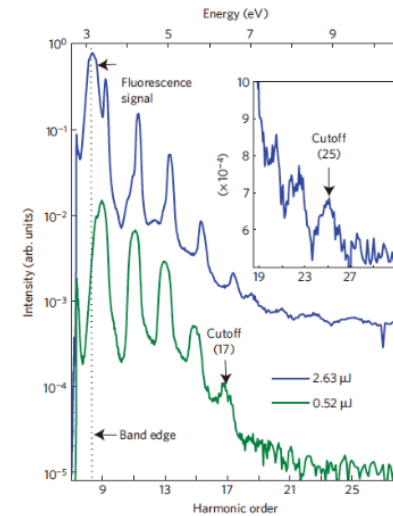
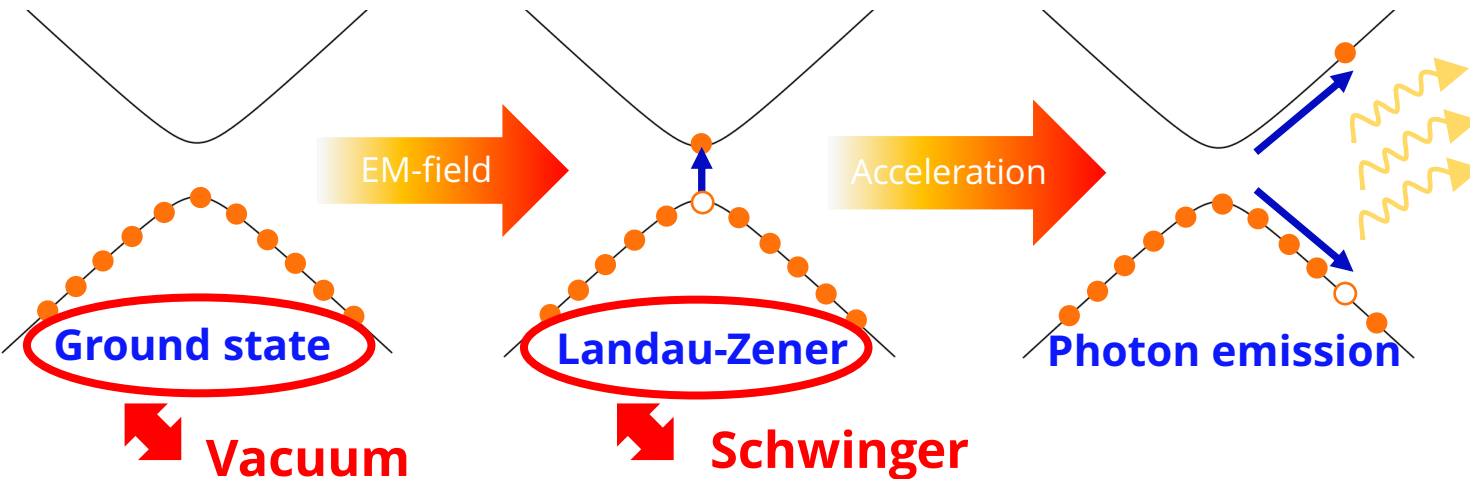
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- Theory for band materials developed recently [Vampa et al. (2014)]
(cf. 3 step model for gas is insufficient for band materials) [Corkum (1993)] [Lewenstein et al. (1994)]
- But still incomplete and remains a hot topic, e.g., lack of analytical formulation

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Analogy b/w Schwinger & Landau-Zener \Rightarrow natural to expect HHG from the vacuum

How to formulate vacuum HHG ?

Observable: Photon spectrum \Rightarrow solving of a differential equation

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\therefore We can take two strategies to discuss vacuum HHG:

(1) Numerically solve Eq. (\spadesuit) (sometimes called TDSE method)

(2) Analytically solve Eq. (\spadesuit) by using some "wise" mathematical method

\Rightarrow In this work: Exact WKB method

Exact WKB method

Exact WKB = a wise method to solve ODE with a small parameter

[Voros (1983)]
[Pham, Dillinger, Delabaere,
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Exact WKB method

Exact WKB = a wise method to solve ODE with a small parameter

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A perturbation theory w.r.t. \hbar (or adiabatic approx.)

• Consider $0 = [\hbar^2 \partial_t^2 + Q(t)]\phi(t) \stackrel{t \equiv \hbar\tau}{\Leftrightarrow} [\partial_\tau^2 + Q(\hbar\tau)]\phi(\tau)$

$$\Rightarrow \phi_{\pm}(t; \hbar) := \exp\left[\mp \frac{i}{\hbar} \int_{t_0}^t dt' \sqrt{Q(t')}\right] \times \sum_{n=0}^{\infty} \psi_{\pm,n}(t) \hbar^n$$

0th order = plane wave
 $\sim \exp[\mp \frac{i}{\hbar} \sqrt{Q}t]$

Perturbation w.r.t. \hbar

- WKB expansion makes sense if the **perturbative part** is convergent
- However, $\psi_{\pm,n} \sim n!$ in general (e.g., Airy function $Q(t) \propto t$)

\Rightarrow WKB expansion has zero radius of convergence \Rightarrow ill-defined !!!

Exact WKB method

Exact WKB = a wise method to solve ODE with a small parameter

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A resummation scheme for factorially divergent $\sim n!$ series

- Consider the div. part of WKB expansion $\psi_{\pm}(t; \hbar) := \sum_{n=0}^{\infty} \psi_{\pm,n}(t) \hbar^n$
 - ① Construct "Borel transformation": $B[\psi_{\pm}](t; \eta) := \sum_{n=0}^{\infty} \frac{\psi_{\pm,n}(t)}{n!} \eta^n$
 - ② Laplace trans. gives "Borel sum": $\Psi_{\pm}(t; \hbar) := \int_0^{\infty} \frac{d\eta}{\hbar} e^{-\eta/\hbar} B[\psi_{\pm}](t; \eta)$
- Ψ_{\pm} is well-defined and is a natural analytic continuation of ψ_{\pm}
 \Rightarrow **Ψ_{\pm} gives a well-defined version of the WKB solution !**
- Note 1: If $B[\psi_{\pm}]$ has singularities on the real axis, Ψ_{\pm} is ill-defined
 \Rightarrow but, the "ill-defined-ness" can be used to quantify the Stokes phenomenon of the WKB solution
- Note 2: in practice, some approximations need to be used in Borel resum
 \Rightarrow in this work: only leading-order $n!$ div. is considered (Airy approx.)

Result (1/2)

Setup: A monochromatic E-field $E(t) = E_0 \cos(\Omega t)$ applied onto QED vacuum

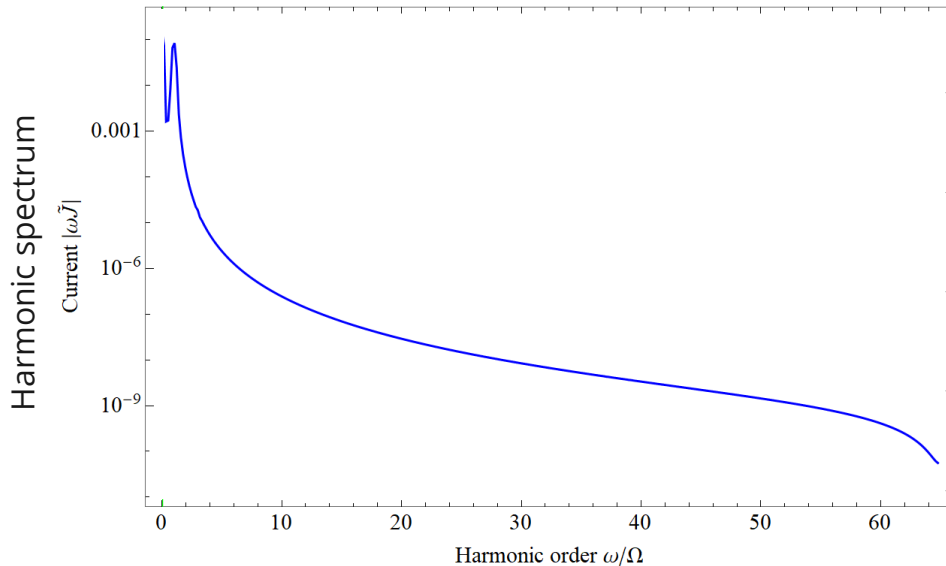
What I did: Compute harmonic spectrum **numerically** and **analytically with WKB**

Result (1/2): Numerical

Setup: A monochromatic E-field $E(t) = E_0 \cos(\Omega t)$ applied onto QED vacuum

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Weak field: $\frac{m\Omega}{eE_0} = 100$

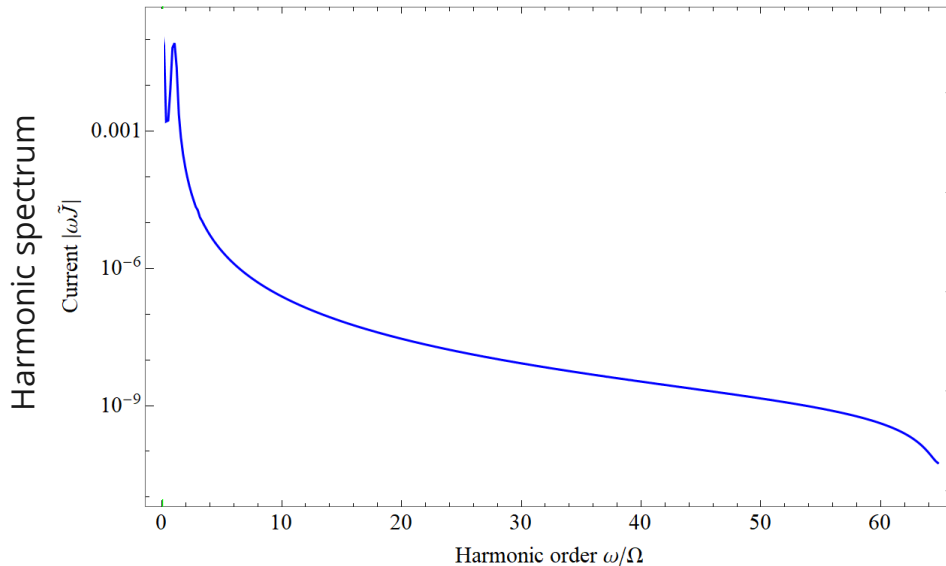


Result (1/2): Numerical

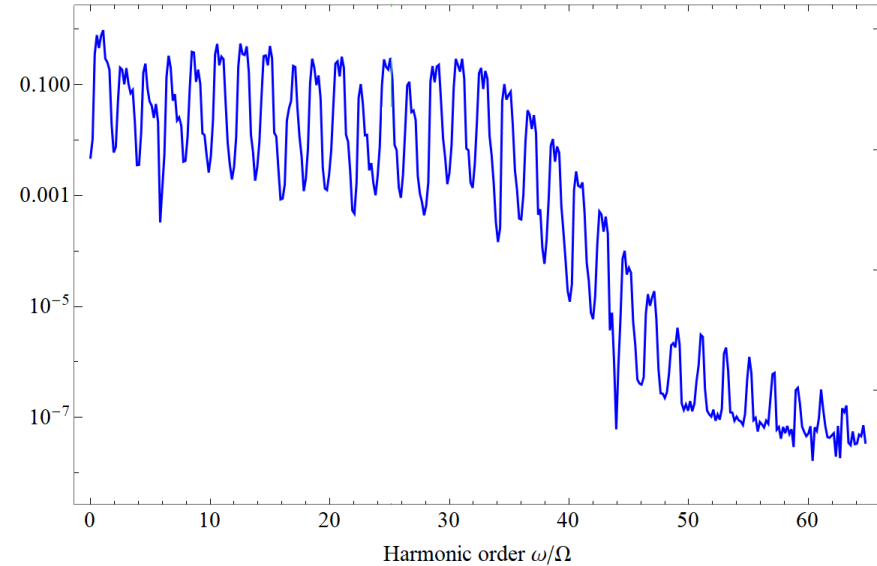
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Weak field: $\frac{m\Omega}{eE_0} = 100$



Strong field: $\frac{m\Omega}{eE_0} = 1/4$

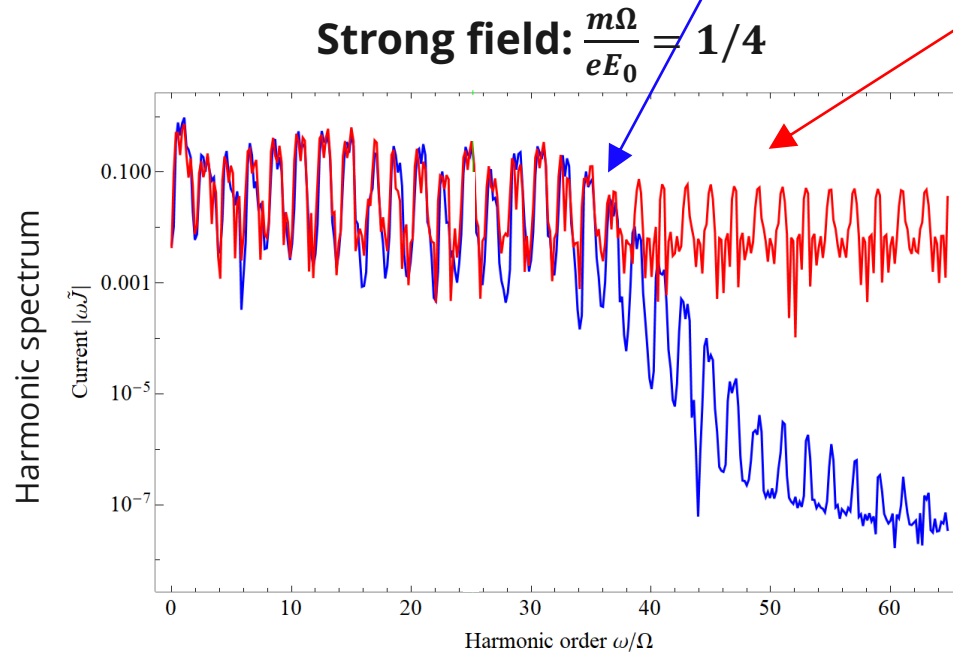


- **High harmonics appears from the vacuum** when field becomes strong !
- structure similar to solid HHG (e.g., plateau, cutoff)

Result (2/2): Analytical

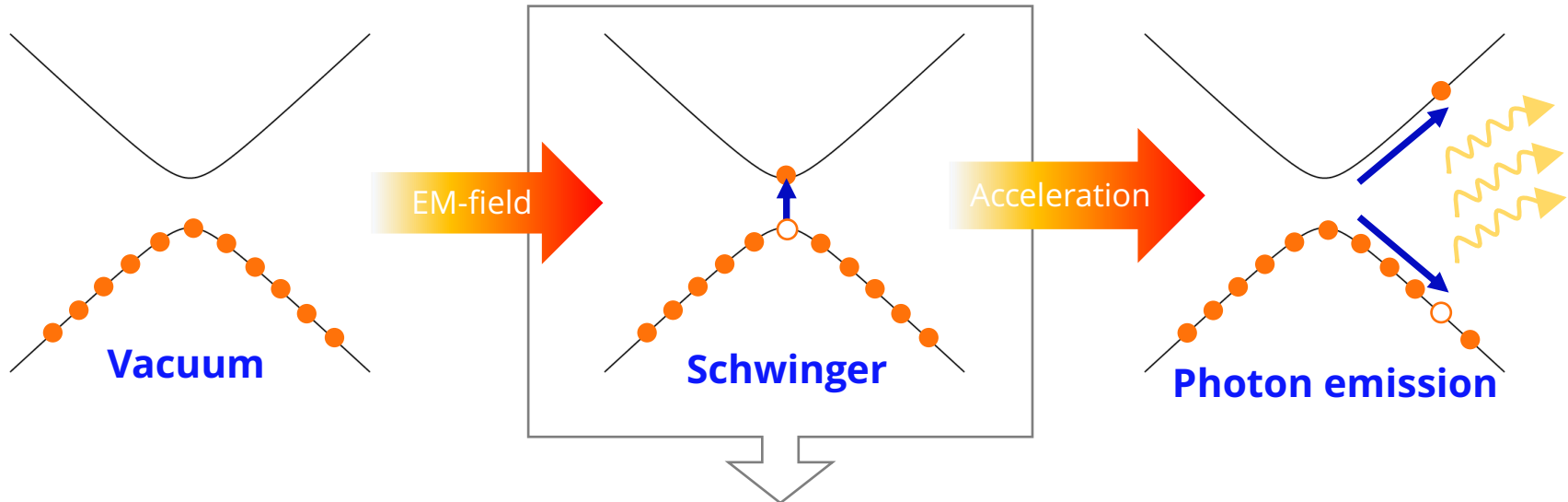
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What I did: Compute harmonic spectrum **numerically** and **analytically with WKB**



- **High harmonics appears from the vacuum** when field becomes strong !
 - structure similar to solid HHG (e.g., plateau, cutoff)
- Analytical WKB calculation is good at the plateau but bad after the cutoff
 - ⇒ useful to get a deeper insight of HHG in the agreement region e.g., determination of the cutoff location (→ next slide)

Physics picture: Interband interference

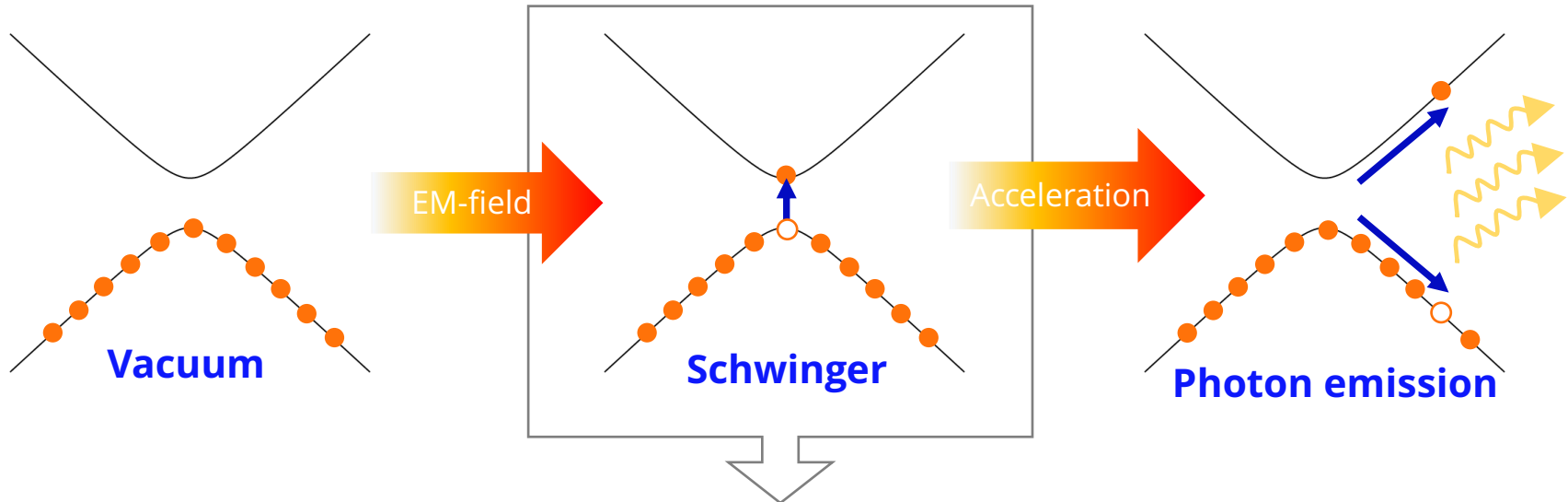


- Total wavefunction = superposition of the two energy states:

$$\frac{eE_0}{\Omega} \cos(\Omega t)$$

$$\psi = \psi_+ + \psi_- = \# e^{-i\int \epsilon dt} + \# e^{+i\int \epsilon dt} \quad \text{w/} \quad \epsilon \sim \sqrt{m^2 + \mathbf{p}^2} \sim \sqrt{m^2 + (e\mathbf{A}(t))^2}$$

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⇒ current can be decomposed into **intra-** and **inter-**band contributions

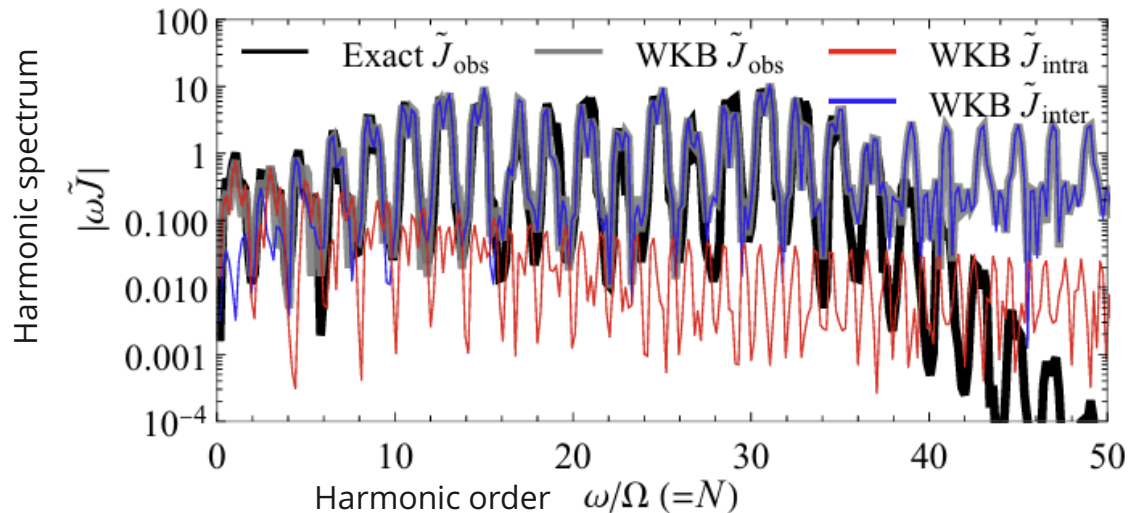
$$J \sim \bar{\psi} \gamma \psi \sim \# \bar{\psi}_{\pm} \psi_{\pm} + \# \bar{\psi}_{\pm} \psi_{\mp} \sim \# e^{i\int (\epsilon - \epsilon) dt} + \# e^{i\int (\epsilon + \epsilon) dt} \sim \# + \# e^{2i\int \epsilon dt}$$

⇒ the intra-band interference has higher frequency ⇒ **source of HHG**

- is different from the naïve 3 step model in gas HHG, but is consistent w/ solid HHG

[Vampa et al. (2014)]

Physics picture: Interband interference



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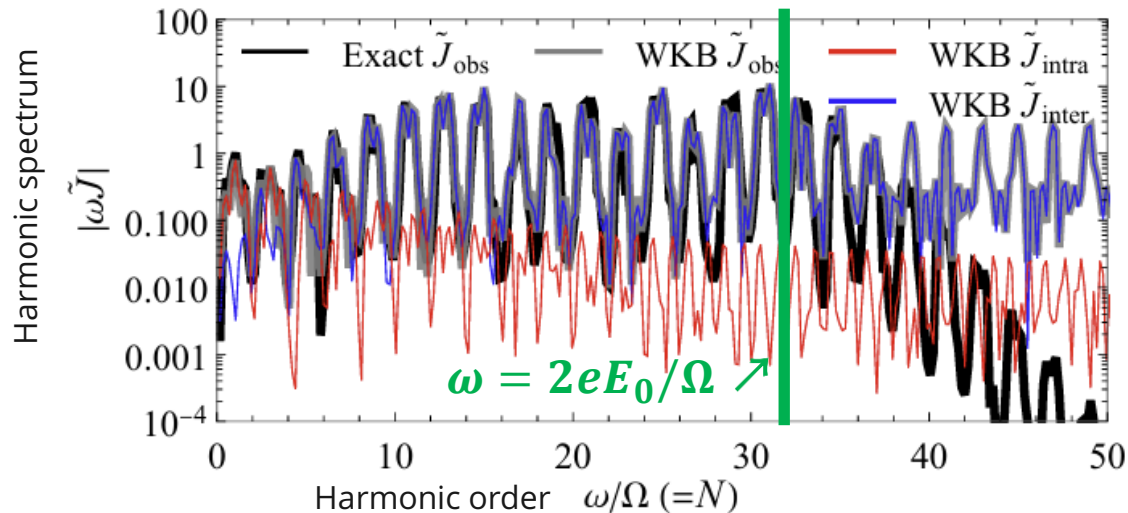
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[Vampa et al. (2014)]

- Cutoff law: F.T. $[J_{\text{inter}}] \sim \int dt \exp[-i \int dt (\omega - 2\epsilon)] \Rightarrow \omega_{\text{cutoff}} \sim \max_t 2\epsilon(t) \sim 2eE_0/\Omega$
 - $\omega_{\text{cutoff}} \propto E_0$ is consistent w/ solid HHG, while $\omega_{\text{cutoff}} \propto \Omega^{-1}$ is our prediction

[Ghimire et al., (2011)]

Contents

1. Overview of strong-field physics

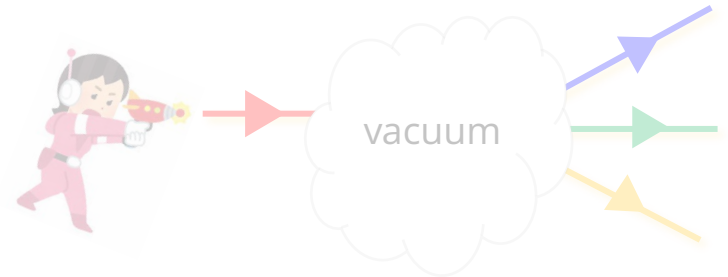
- becoming a hot topic in physics, thanks to the recent developments in exp.
- the Schwinger effect = a process to create “something” from “nothing”

2. Application of the Schwinger effect

[HT, Hongo, Ikeda (2021)]

⇒ High-harmonic generation from the vacuum

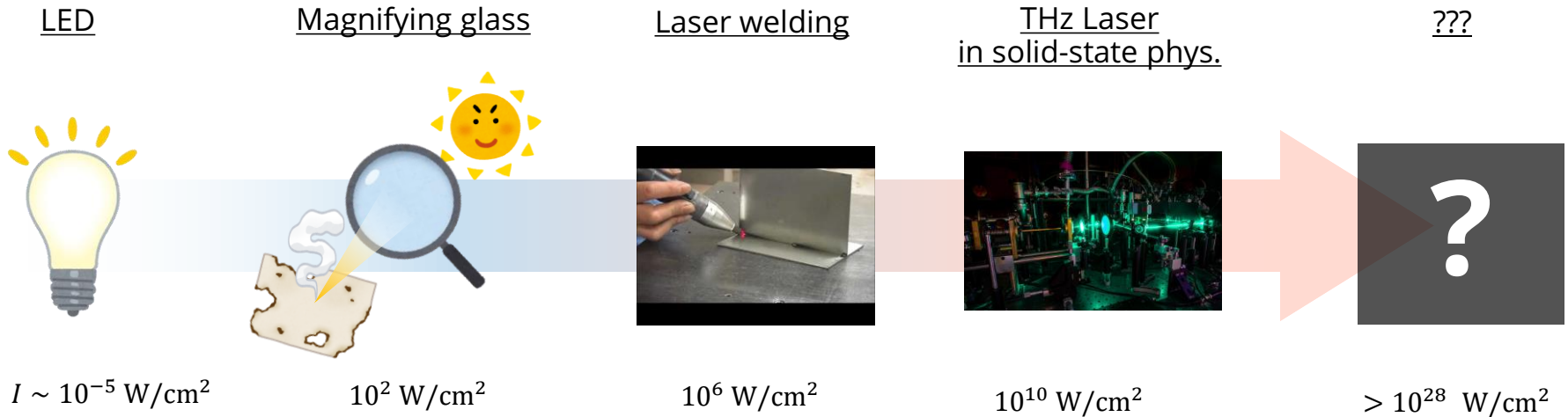
- first prediction of HHG from the vacuum
- a good example of “interdisciplinarity” of strong-field physics
- ⇒ your expertise should be useful in my area and vice versa !



3. Summary

Summary

Q: What happens if we make light (or “field” in general) stronger and stronger ?

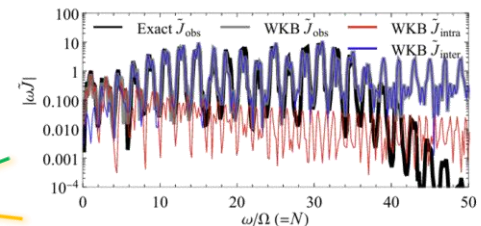
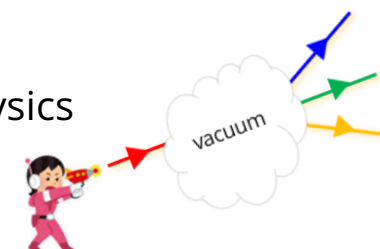


A: Our vacuum shall be destroyed and gives us nontrivial responses

- Why interesting: Unexplored nonpert. regime \Rightarrow many nontrivial phenomena we've never seen
- Why timely: Recent availability of strong fields
e.g., high-power laser, magnetar, heavy-ion colls

Example phenomenon: HHG from the vacuum

- First prediction of HHG from the vacuum
- An example of “interdisciplinarity” strong-field physics
 \Leftarrow borrowed ideas from optics, solid-state physics, & math





Exact-WKB recipe for Stokes phenomenon

Step 1: Draw a Stokes graph

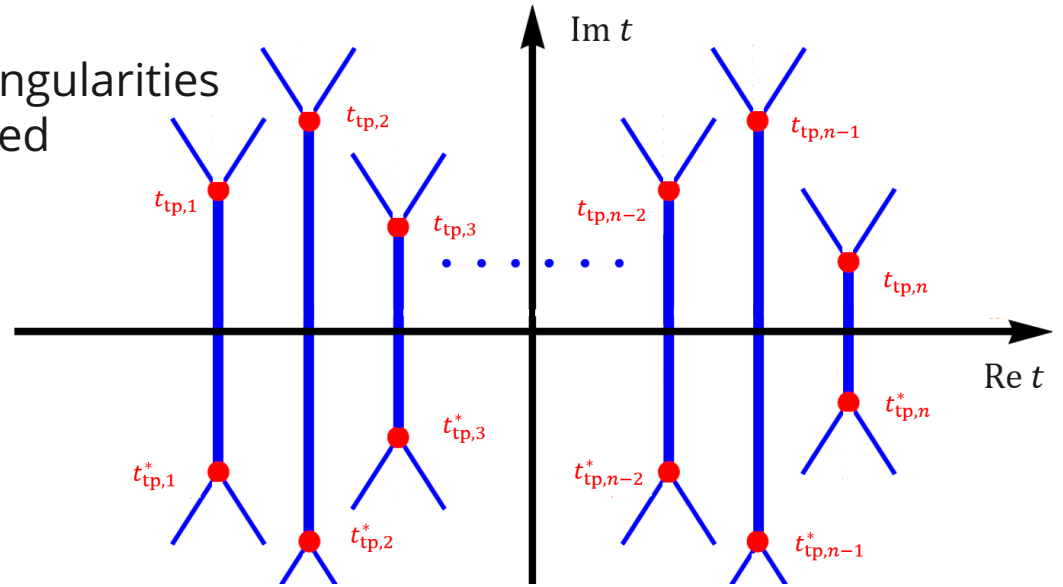
- Laplace trans. of $B[\psi_{\pm}](t; \eta)$ hits singularities (non-Borel summable) when located on **Stokes lines** in the t -plane

— : Stokes lines

$$\{t \in \mathbb{C} \mid 0 = \text{Im} i \int_{t_{\text{tp}}}^t dz \sqrt{Q(z)}\}$$

● : turning points

$$Q(t_{\text{tp}}) = 0$$



Step 2: Compute Borel sum Ψ_{\pm} at each Stokes region

- Borel sum is well-defined and computable in each region separated by Stokes lines

$$\Psi_{\pm} = \int_0^{\infty} \frac{d\eta}{\hbar} e^{-\eta/\hbar} B[\psi_{\pm}](t; \eta) \sim \exp\left[\mp \frac{i}{\hbar} \int_{t_0}^t dt' \sqrt{Q(t')}\right] \times (1 + O(\hbar)) \text{ at each Stokes region}$$

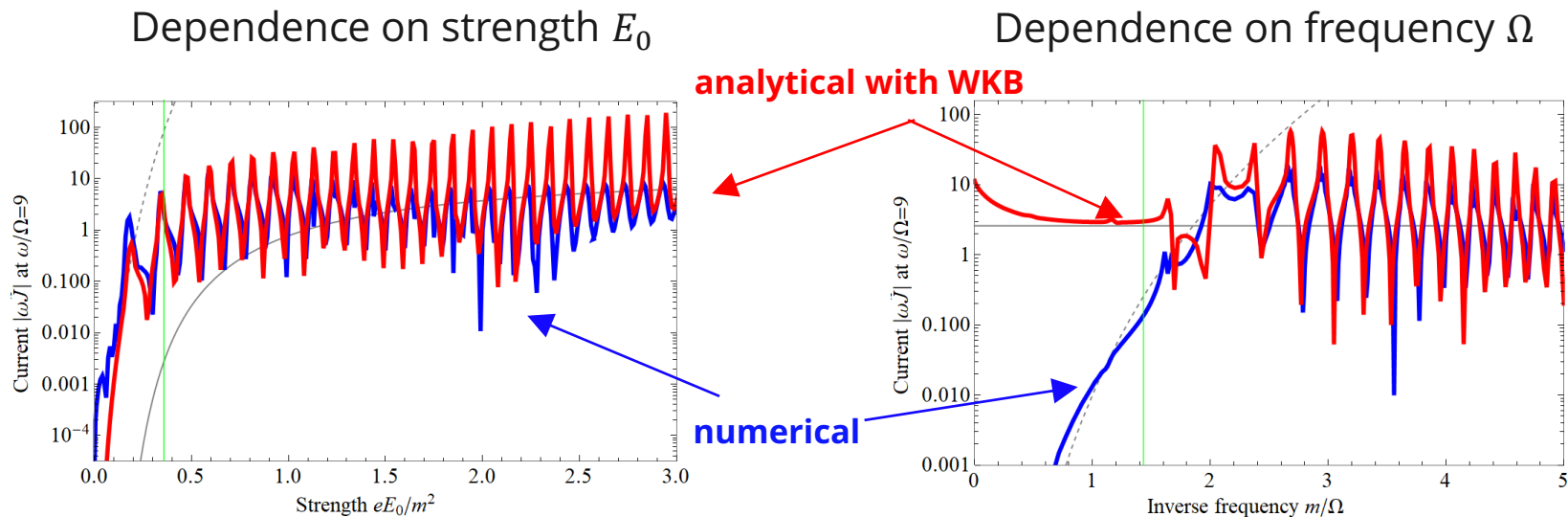
Step 3: Compute the Stokes constants α and β

- Whenever crosses Stokes lines, Ψ_{\pm} jumps discontinuously (**Stokes phenomenon**)
 \Rightarrow The discontinuity is given by the integral of singularities of $B[\psi_{\pm}]$

$$\Psi_+(\text{region A}) = \alpha \Psi_+(\text{region B}) + \beta \Psi_-(\text{region B}) \sim \oint_{\text{sing. on a Stokes line}} \frac{d\eta}{\hbar} e^{-\eta/\hbar} B[\psi_{\pm}](t; \eta)$$

Application to high-harmonic generation (2/2)

- WKB works more in the deep non-perturbative regime $E_0 \rightarrow$ large, $\Omega \rightarrow$ small
- Demonstration: magnitude of the harmonic peak at $\omega/\Omega = 9$



• Lessons:

- (1) WKB makes it easier to analyze the non-perturbative regime
- (2) Saturation & oscillation of the harmonic intensity

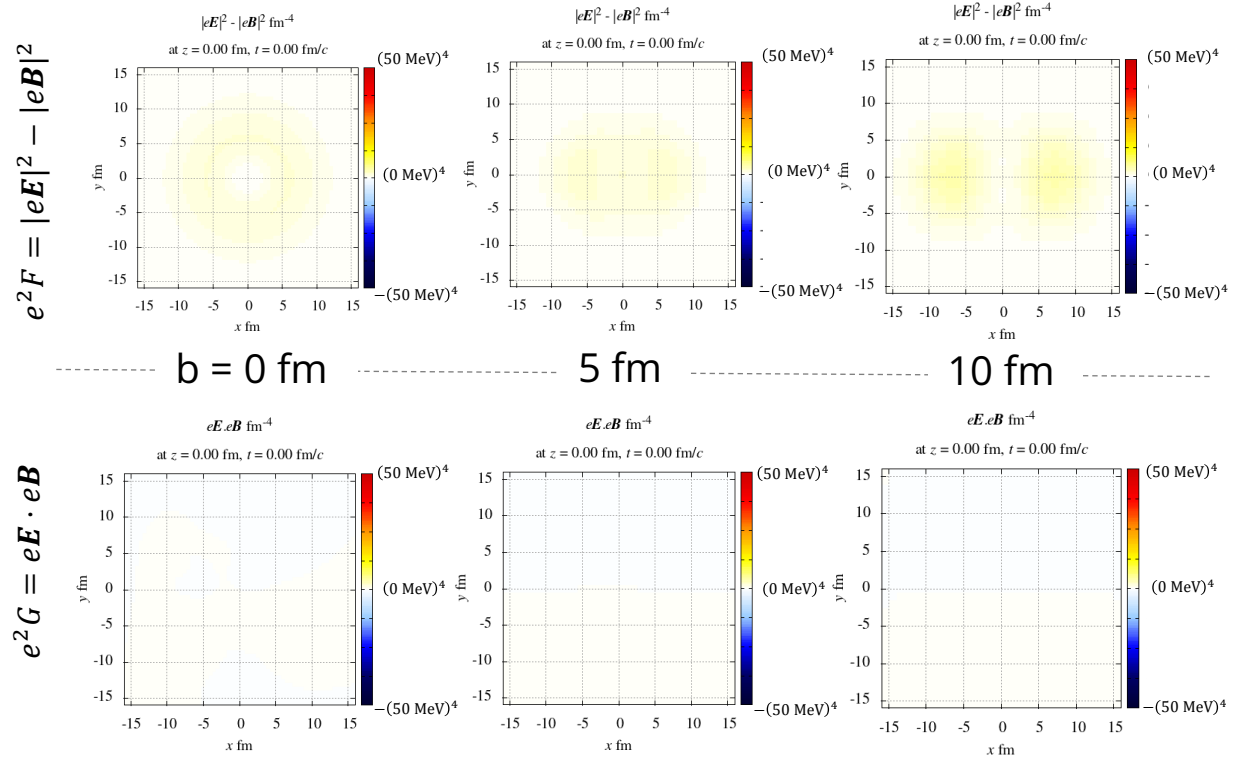
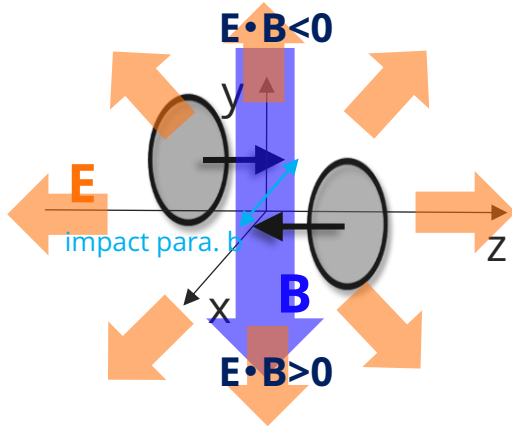
\Rightarrow consistent with recent semi-conductor exp. [Xia et al., (2020)]

(but only E_0 -dep. is measured and Ω -dep. is our prediction)

Peripheral case

[Nishimura (Osaka), Ohnishi (Kyoto), HT, in progress]

✓ Non-central collision \Rightarrow \mathbf{B} & $\mathbf{E} \cdot \mathbf{B}$ are produced



Preliminary results

@ $E_{CM} = 7.7$ GeV

- $e\mathbf{E} \cdot e\mathbf{B} = O((50 \text{ MeV})^4) \Rightarrow$ Non-negligible for QCD (Enough strong for electron and current quarks)
- Perhaps, a nice place to study chiral-anomaly-related stuffs

ex) chirality production via the Schwinger mechanism
$$N_5 = VT \frac{e\mathbf{E} \cdot e\mathbf{B}}{2\pi^2} \exp\left[-\pi \frac{m^2}{eE}\right]$$